

# Classical Phase Space Density for the Relativistic Hydrogen Atom

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**Abstract.** Quantum mechanics is considered to arise from an underlying classical structure (“hidden variable theory”, “sub-quantum mechanics”), where quantum fluctuations follow from a physical noise mechanism. The stability of the hydrogen ground state can then arise from a balance between Lorentz damping and energy absorption from the noise. Since the damping is weak, the ground state phase space density should predominantly be a function of the conserved quantities, energy and angular momentum.

A candidate for this phase space density is constructed for ground state of the relativistic hydrogen problem of a spinless particle. The first excited states and their spherical harmonics are also considered in this framework.

The analytic expression of the ground state energy can be reproduced, provided averages of certain products are replaced by products of averages. This analysis puts forward that quantum mechanics may arise from an underlying classical level as a slow variable theory, where each new quantum operator relates to a new, well separated time interval.

## INTRODUCTION

In this centenary of Einstein’s *Anns Mirabilis* it is appropriate to reconsider the foundations of physical theories. We shall, in particular, be interested in a possible bridge between the three fields that Einstein pushed so much in 1905: brownian motion, special relativity and quantum mechanics. Our focus is to investigate the possibility that quantum mechanics arises as a statistical theory from a certain special relativistic theory, with specific stochastic forces constituting the notorious “quantum noise”. Conceptually this will be a “unified field theory”, which Einstein tried to formulate in the second part of his life. Without fully specifying this theory, we shall generalize squares of quantum wavefunctions to phase space densities and see what it leads to.

In order to motivate our investigations, let us notice that the status of quantum mechanics is still judged in two basically opposite ways. The empirical view stresses that quantum mechanics, with its many postulates such as the superposition principle, the uncertainty principle, the exclusion principle, the Born rule, the collapse in measurements, is our best theory of nature, proven to work in all cases where it was tested, ranging from solid state physics and quantum chemistry to high energy physics and the early Universe. Most physicists even believe that quantum mechanics is the ultimate theory of nature, that must be unified with gravity, one popular route being string theory, which adopts the additional postulates of supersymmetry and extra dimensions.

On a parallel track, there has always remained the fundamental question: what is quantum mechanics actually standing for? Does nature have a reality: do particles exist as physical entities, possessing a position and a momentum; is the moon there when nobody looks? What is the physical mechanism behind the notorious “quantum fluctuations” and uncertainty relations? What causes interference effects and the particle-wave duality? Does each particle have its own wavefunction (standard Copenhagen interpretation), what does a “collapse of the wavefunction” mean physically<sup>1</sup>? For decades these questions have suffered from the lack of progress, in particular in the quantum measurement problem, the only undisputed point of contact between quantum mechanics and nature. Very different interpretations have emerged [2]: the Everett-Wheeler relative state or multi-universe picture (collapse does not occur, but rather a branching); the mind-body problem (the registration of the measurement in our mind or a machine matters) (Wigner, see [2]), the wavefunction codes our state of knowledge [3] or our state of belief [4].

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<sup>1</sup> Nothing measurable according to Zbinden et al [1].

At present a variety of different approaches support the possibility that at the quantum level, nature has a ‘natural’ (i.e. more classical than long supposed) behavior. We may mention: arguments against the conclusion that nature should be non-local because of violations of Bell inequalities [5, 6, 7]; the demonstration that quantum probabilities including interference effects are compatible with classical probability theory [8]; the long effort of Stochastic Electrodynamics [9] and its later version called Linear Stochastic Electrodynamics [9, 10]; the demonstration of entanglement in classical brownian motion [11]; the description of an electron with its spin as a soliton (over-extremal Kerr-Newman black hole) in electro-gravity [12, 13]; the many arguments in favor of the statistical interpretation [14, 15, 16, 17, 18, 19]; explicit solutions for the quantum measurement problem [18, 19].

Let us go in some detail about the measurement process. By taking into account an explicit model for the apparatus, the quantum measurement problem was recently solved in our group in two situations: in the first the position of a boson is measured by an apparatus that consists of many bosons. The apparatus starts close to an (ideal) Bose-Einstein phase transition and is driven into it by the measurement [18]. In the second model the  $z$ -component of a spin- $\frac{1}{2}$  is measured by an apparatus consisting of many spins- $\frac{1}{2}$  [19] that start out as a paramagnet and are driven into the up or down ferromagnetic phase, according to the sign of the measured spin component. In both cases the apparatus also consists of a heat bath, which originates from other degrees of freedom, in the magnetic case the lattice vibrations.

Quantum mechanics being a probabilistic theory, it should come as no surprise that what could be calculated was the statistical distribution of outcomes for an ensemble of measurements on an ensemble of identically prepared systems or identical preparations of a single system. Technically this comes in because the whole exercise is to determine the post-measurement density matrix given its pre-measurement form. The process goes in two steps: when coupling the system to the apparatus, there appears a very fast collapse of the off-diagonal terms (the death of Schrödinger cats). This “collapse of the wavepacket” is a unitary evolution in the large Hilbert space of system and pointer degrees of freedom of the apparatus (thus not involving the bath), very similar to the process of dephasing in NMR physics. The collapse time depends as an inverse power on the number of particles in the apparatus, so it is strongly microscopic when the apparatus is macroscopic. On a later stage the effect of the bath is to make this collapse definite by decoherence. Schrödinger cat terms thus disappear automatically, because of the macroscopic size of the apparatus. The second step of the measurement is its registration, which involves the diagonal elements of the density matrix and appears to have classical features [20]. Indeed, for the magnetic model, the solution of this part of the problem is just what one might have anticipated from purely classical thermodynamic reasoning [20]. In particular, a bath is needed simply because energy has to be dumped when the macroscopic pointer variable (the magnetization of the apparatus) goes from its initial metastable paramagnetic value to its up or down ferromagnetic value. Also in cloud chambers, the energy released upon creating bubbles in an oversaturated liquid has to be transferred to the environment and this very mechanism allows the selection of circumstances where these bubbles have a desired average size.

From the recent solutions of two realistic models for the quantum measurement process, it became clear that the statistical interpretation of quantum mechanics allows a simple interpretation of this process during and after the measurement. To mention one problem that is immediately solved in this way: the fundamental question how quantum mechanics can describe collapse of the measured operator has the answer: quantum mechanics describes only the statistics of those collapses; it is not the proper theory to describe an individual collapse. Spontaneous collapse models (GRWP theories), that try to treat individual collapses from a slightly extended quantum mechanics, attempt to heal a problem that has its roots at a more fundamental level. Indeed, for a two slit experiment quantum mechanics does not describe any individual spot on the photographic plate, but only the density of many spots.

Thus, we conclude that *Quantum mechanics is not a theory of nature itself, but a theory for the statistics of outcomes of experiments*. This situation is very similar to the one of thermodynamics, where the underlying level of, say a Lennard-Jones gas, is Newton’s equations of motion. The main difference is, however, that experiments have to be made with apparati which themselves are always macroscopic and therefore influence the microscopic dynamics. Since it is the task of physicists to describe nature, the search for a sub-quantum theory (hidden-variables theory) is, in our eyes, completely justified. The program to search this may contain the following steps

- 1) For point particles, show compatibility between classical phase space density structures and quantum structures.
- 2) Find the underlying stochastic theory and isolate the stochastic forces which cause the “quantum fluctuations”. Show compatibility with quantum mechanics.
- 3) Find soliton-like solutions for elementary particles such as the electron. An interesting attempt sees an electron as a Kerr-Newmann black hole, that is, as a charged ring with diameter given by the Compton wavelength [12, 13]. On the ring a current runs with the speed of light, causing a spin, with  $g = 2$ , as in the Dirac theory for the electron. In these theories the parameters must be fixed to the physical values by adding terms to the Einstein-Maxwell theory. A possibility is to couple torsion with the electromagnetic field tensor [21], which might fix charge and/or spin, though

not yet the mass. Within this framework the Pauli principle might arise as an energetic constraint on the behavior in the far field region, and Einstein photons as massless spin-one solitons. Let us notice that indistinguishably is a natural aspect of solitons that wind up in a chaotic motion.

4) Show that the consistency still works for the “stochastic soliton mechanics” of solitons in the random electromagnetic background, with “photons” appearing as classical EM waves and as “Einstein” solitons.

Clearly, the program may run into contradictions at many points: apart from internal consistency, compatibility with a huge amount of experiments is required. But if Einstein was right after all, then there should be some road to the underlying structure, and this one seems to us the simplest, so it deserves to be investigated further.

In this work we shall restrict ourselves to point 1) of the program and observe that already this one is more complicated than sketched here.

## MOTIVATION FROM STOCHASTIC ELECTRODYNAMICS

Quantum mechanics has both a particle and wave nature. In a subquantum approach, it is clear how to think of the particles, to start they are just considered as points, so the next question: what are the waves made of. In the simplest approach, one assumes that the waves are due to fields that we know already. The first candidate then is Electromagnetism. One has then arrived at the so-called Stochastic Electrodynamics (SED), where the zero-point fluctuations known as a slogan in quantum mechanics are attributed to a physical entity, namely classical random EM fields [9]. To get the proper connection, one then has to choose the energy content of each mode as  $\frac{1}{2}\hbar\omega$ , in which the prefactor  $\hbar$  is just a system parameter, chosen to coincide with Planck’s constant. The related spectral density is just the zero point part of the Planck spectrum,  $\rho(\omega) = \hbar\omega^3/(2\pi^2c^3)$ . This  $\omega^3$ -dependence is precisely the one consistent with Lorentz invariance and with thermodynamics.

The theory of SED has a long history of both success and failures. SED has explained the harmonic oscillator and, with it, electric dipole harmonic oscillators, the Casimir and Casimir/Polder effects, and the Unruh effect.

Away from harmonic systems, SED has a long record of failures. First, a connection with the Schrödinger equation has never been found. Second, the role for excited states has been discussed only in the margin. It should be admitted that, whereas Schrödinger dynamics at first neglects the decay of excited states, the structure of SED is that of time-dependent perturbation theory, where the full time-dependence is taken into account from the start. Third, a long list of non-linear problems has been studied, where results different from quantum mechanics emerge, see e.g. [22].

The most obvious example will be the subject of this study: the ground state of the hydrogen problem. This is a rather clean problem that should be understood before conclusions are drawn about validity or invalidity and about multi-level atoms. Indeed, in the hydrogen problem no internal structure should be expected that complicates the orbiting of the point-electron in the central force field of the infinitely heavy nucleus. Theoretical results so far are meager, predicting stability radius of the orbit equal to infinity [9]. Surprisingly, a numerical analysis by Zhou and Cole supports an SED explanation of the quantum ground state density [23].

That the hydrogen problem should have a stable solution within SED, was put forward long ago: when the electron is far from the nucleus, its orbital frequency is low and so should be the relevant frequencies of the stochastic force. Due to the  $\omega^3$  spectrum, there should essentially be no random force at all. What remains is the Lorentz damping that brings the electron closer to the nucleus. For small distances, on the other hand, the relevant frequencies should be large, and the spectrum predicts a strong kicking due to the random forces, preventing electron to fall into the nucleus due to Lorentz damping alone. There thus appears a stable atomic state. Is this the one we know from quantum mechanics? So far attempts to show this have failed, including those by ourselves [21]. For this reason it would be interesting to corroborate the numerical analysis of [23].

## QUANTUM MECHANICS OF THE RELATIVISTIC HYDROGEN ATOM

The Hamiltonian for a spinless boson with rest mass  $m$  and charge  $e$ , in the presence a central charge  $-Ze$ , reads

$$H = \sqrt{m^2c^4 + \mathbf{p}^2c^2} - \frac{Ze^2}{r} = \gamma mc^2 - \frac{Ze^2}{r} \quad (1.1)$$

We express  $\mathbf{r}$ ,  $\mathbf{p}$  and time in atomic units,

$$\frac{\mathbf{r}}{a_0} \rightarrow \mathbf{r}, \quad \frac{t}{\tau_0} \rightarrow t, \quad \frac{\mathbf{p}\tau_0}{ma_0} \rightarrow \mathbf{p} \quad (1.2)$$

with Bohr radius and Bohr time

$$a_0 = \frac{\hbar}{\alpha Z mc} = \frac{1}{Z} 5.29 \cdot 10^{-11} m, \quad \tau_0 = \frac{\hbar}{\alpha^2 Z^2 mc^2} = \frac{1}{Z^2} 2.418 \cdot 10^{-17} s \quad (1.3)$$

with numerical values for the electron. This scaling also implies

$$\frac{E}{mc^2} \rightarrow E, \quad \frac{L}{\hbar} \rightarrow L, \quad (1.4)$$

and brings the scaled form

$$H = \sqrt{1 + \alpha^2 Z^2 \mathbf{p}^2} - \frac{\alpha^2 Z^2}{r} = \gamma - \frac{\alpha^2 Z^2}{r} \quad (1.5)$$

Let us denote quantum operators by a subscript  $op$ . The Schrödinger equation reads

$$(H_{op} - E)|\psi\rangle = (\sqrt{1 + \alpha^2 Z^2 \mathbf{p}_{op}^2} - \frac{\alpha^2 Z^2}{r_{op}} - E)|\psi\rangle = 0 \quad (1.6)$$

this is equivalent to

$$(\frac{1}{2}\mathbf{p}_{op}^2 - \frac{\alpha^2 Z^2}{2r_{op}^2} - \frac{E}{r_{op}})|\psi\rangle = -\frac{1-E^2}{2\alpha^2 Z^2}|\psi\rangle \quad (1.7)$$

In the position basis this reads

$$(-\frac{1}{2}\partial_r^2 - \frac{1}{r}\partial_r + \frac{L_{op}^2 - \alpha^2 Z^2}{2r^2} - \frac{E}{r} + \frac{E^2}{R})\psi(r) = 0, \quad \frac{1}{R} = \frac{1-E^2}{2\alpha^2 Z^2 E^2} \quad (1.8)$$

When scaling  $r \rightarrow \tilde{r}/E$ ,  $\mathbf{p} \rightarrow E\tilde{\mathbf{p}}$ , and using that  $L_{op}^2$  has eigenvalues  $l(l+1)$  with non-negative integer  $l$ , this becomes

$$(-\frac{1}{2}\partial_{\tilde{r}}^2 - \frac{1}{\tilde{r}}\partial_{\tilde{r}} + \frac{l(l+1) - \alpha^2 Z^2}{2\tilde{r}^2} - \frac{1}{\tilde{r}})\psi = -\frac{1}{R}\psi \quad (1.9)$$

so  $-1/R$  is the associated non-relativistic energy. It is customary define the effective angular momentum value  $\ell_l$  by

$$\ell_l(\ell_l + 1) = l(l+1) - \alpha^2 Z^2, \quad \ell_l = -\frac{1}{2} + \frac{1}{2}\sqrt{(2l+1)^2 - 4\alpha^2 Z^2} \approx l - \frac{\alpha^2 Z^2}{2l+1} \quad (1.10)$$

Now the whole problem can be solved by analogy with the non-relativistic situation. The ground state wavefunction is

$$R_{10} = Cr^{\ell_0} e^{-rE_{10}/(\ell_0+1)}, \quad (1.11)$$

The so called Yrast states have  $(n, l) = (n, n-1)$  and wavefunctions

$$R_{nl} = Cr^{\ell_l} e^{-rE_{nl}/(\ell_l+1)} \quad (1.12)$$

They have energy

$$E_{n,l} = \left\{ 1 + \frac{\alpha^2 Z^2}{[n - l - \frac{1}{2} + \sqrt{(l + \frac{1}{2})^2 - \alpha^2 Z^2}]^2} \right\}^{-1/2} \approx 1 - \frac{\alpha^2 Z^2}{2n^2} - \alpha^4 Z^4 \left[ \frac{1}{n^3(2l+1)} - \frac{3}{8n^4} \right] \quad (1.13)$$

a formula which is valid for all levels, that is to say, for  $n = 1, 2, \dots$  and  $0 \leq l \leq n-1$ .

Let us mention that when the spin of the electron is taken into account, one has to replace  $l \rightarrow j = l \pm \frac{1}{2}$ .

## STATISTICAL MECHANICS OF THE RELATIVISTIC COULOMB PROBLEM

The Schrödinger equation works so well for the hydrogen atom, because corrections arising from the coupling to the electromagnetic field are weak. They lead to the Lamb shift, which is of order  $\alpha^5 Z^4 \ln \alpha Z$  in the present units. So it is weaker than the fine structure, represented by the  $\alpha^4 Z^4$  term in (1.13). The very same mechanism makes the lifetimes of excited states much larger than the Bohr time.

To focus our line of reasoning, let us consider the theory of stochastic electrodynamics (SED). Here random electromagnetic fields are assumed to exist. In this theory a Fokker-Planck approach may be formulated for the statistics of the random orbits  $(\underline{\mathbf{r}}(t), \underline{\mathbf{p}}(t))$  [9, 21]. The phase space density will satisfy an equation of the form

$$P(\mathbf{r}, \mathbf{p}, t) = \langle \delta(\underline{\mathbf{r}}(t) - \mathbf{r}) \delta(\underline{\mathbf{p}}(t) - \mathbf{p}) \rangle \quad (2.1)$$

will satisfy

$$\partial_t P = -\mathcal{L}P + \text{Lorentz damping} + \text{diffusion terms} \quad (2.2)$$

with drift given by the Liouvillian

$$\mathcal{L} \equiv \frac{1}{\gamma} \mathbf{p} \cdot \nabla_{\mathbf{r}} - \frac{\mathbf{r}}{r^3} \cdot \nabla_{\mathbf{p}} \quad (2.3)$$

Since damping and diffusion terms are of order  $\alpha^3$ , the stationary distribution should, to this order, be a function of the conserved quantities, angular momentum  $L$  and energy  $E$ . This remains valid if one goes beyond the Fokker-Planck equation and for other stochastic theories. So the weakness of the Lamb shift finally allows us to consider equilibrium structures of a class of stochastic theories, without going to specific details.

The ground state should follow from the dynamical problem  $H \equiv \gamma - \alpha^2 Z^2/r = E$  or, equivalently,

$$\frac{1}{2} \mathbf{p}^2 - \frac{\alpha^2 Z^2}{2r^2} - \frac{E}{r} = \frac{E^2 - 1}{2\alpha^2 Z^2} \quad (2.4)$$

There will also be effects of damping and noise, but they enter here only at order  $\alpha^3$ , and should explain the Lamb shift. We shall neglect those effects in the present paper. If we define  $R$  by

$$E = \sqrt{1 + \frac{\alpha^4 Z^4}{R^2}} - \frac{\alpha^2 Z^2}{R} = \left(1 + \frac{2\alpha^2 Z^2}{RE}\right)^{-1/2} \quad (2.5)$$

the dynamical problem becomes

$$\frac{1}{2} \mathbf{p}^2 - \frac{\alpha^2 Z^2}{2r^2} - \frac{E}{r} = -\frac{E}{R} \quad (2.6)$$

Close to  $r = 0$  and for  $p_r = 0$  it is seen that  $p_{\perp} = \alpha Z/r$ , implying that  $L \geq \alpha Z$  and  $R \geq r$ . Let us denote

$$\omega = \sqrt{1 - \frac{Z^2 \alpha^2}{L^2}}, \quad \omega^2 L^2 = L^2 - \alpha^2 Z^2 \quad (2.7)$$

We shall take spherical coordinates in  $\mathbf{r}$  space. In the frame along  $\mathbf{r} = r(0, 0, 1)$  we introduce for  $\mathbf{p}$  the angles  $\mu$  and  $\nu$  with  $0 \leq \mu \leq \pi$  and  $0 \leq \nu \leq 2\pi$ ,

$$\mathbf{p} = (p_{\perp 1}, p_{\perp 2}, p_r) = p(\sin \mu \cos \nu, \sin \mu \sin \nu, \cos \mu). \quad (2.8)$$

In special relativity the invariant volume elements in coordinate and momentum space are

$$dV_r = d^3 r = dr d\theta d\phi r^2 \sin \theta, \quad dV_p = \frac{d^3 p}{\gamma} \quad (2.9)$$

We have, using that  $L dL = \omega L d(\omega L)$ ,

$$dV_p = \frac{dp_r dv dp_{\perp} p_{\perp}}{\gamma} = \frac{dp_r dv L dL}{r^2 \gamma} = \frac{dp_r dv \omega L d(\omega L)}{r^2 \gamma} \quad (2.10)$$

We go to the new variables  $\mu$  and  $R$  by the transformation

$$\omega L = r \sqrt{2E(\frac{1}{r} - \frac{1}{R})} \sin \mu, \quad p_r = \sqrt{2E(\frac{1}{r} - \frac{1}{R})} \cos \mu \quad (2.11)$$

and have

$$dp_r d(\omega L) = rd\mu dR \frac{\partial}{\partial R} \left( \frac{E}{r} - \frac{E}{R} \right) = \frac{rd\mu dR \gamma E}{R^2 \sqrt{1 + \alpha^4 Z^4 / R^2}}, \quad \gamma = E + \frac{\alpha^2 Z^2}{r} \quad (2.12)$$

This implies the volume element in momentum space

$$dV_p = \frac{d\mu dv dR}{2R^2 \Phi(E)} \sqrt{\frac{2}{Er} - \frac{2}{ER}} \sin \mu \quad \Phi(E) = \frac{\sqrt{1 + \alpha^4 Z^4 / R^2}}{2E^2} \quad (2.13)$$

Notice that the  $r$ -dependence through  $\gamma$  has dropped out and that  $\Phi = \frac{1}{2}$  non-relativistically. If the phase space density has the form  $\mathcal{P}(E, L) = \omega L R^3 \Phi(E) \exp(-aR)$ , then we have

$$dV_p \mathcal{P} = d\mu dv dR (R - r) e^{-aR} \sin^2 \mu \quad (2.14)$$

The momentum integrals, that is to say, the integrals over  $\mu$ ,  $v$  and  $R \geq r$ , then yield

$$\int_{\mathbf{p}} dV_p \mathcal{P} = \frac{\pi^2}{a^2} e^{-ar} \quad (2.15)$$

which for  $a = 2$  is just the square of the non-relativistic ground state wavefunction. Furthermore, since

$$\frac{\omega^2 L^2 R}{2E} = r(R - r) \sin^2 \mu, \quad (2.16)$$

powers of  $r$  in the squared wave functions can be generated by powers of this ratio in the phase space density. So we may choose the following combination of the conserved quantities  $E$ ,  $R(E)$  and  $L$  for ground state  $(1, 0)$  and, more generally, the Yrast states  $(n, l) = (n, n-1)$  with  $n = 1, 2, \dots$

$$\mathcal{P}(R, L) = C \omega L R^3 \Phi(E) \left( \frac{\omega^2 L^2 R}{2E} \right)^{2\ell_l} e^{-2RE_{nl}/(\ell_l+1)} \quad (2.17)$$

Fixing  $C$  this leads to our final proposition

$$dV_p \mathcal{P} = dR \frac{d\mu}{\pi} \frac{dv}{2\pi} \frac{2^{6+8\ell_l} E_{nl}^{5+4\ell_l}}{(1+\ell_l)^{6+4\ell_l} \Gamma(3+4\ell_l)} r^{2\ell_l} (R - r)^{1+2\ell_l} (\sin \mu)^{2+4\ell_l} e^{-2RE_{nl}/(1+\ell_l)} \quad (2.18)$$

This form is very appealing, though perhaps not unique. The angular integrals yield

$$\int_{\mu, v} dV_p \mathcal{P} = \frac{dR 2^{4+4\ell_l} E_{nl}^{5+4\ell_l}}{(1+\ell_l)^{6+4\ell_l} \Gamma^2(2+2\ell_l)} r^{2\ell_l} (R - r)^{1+2\ell_l} e^{-2RE_{nl}/(1+\ell_l)} \quad (2.19)$$

and the full momentum average brings the proper squares of the wavefunctions,

$$\int_{\mathbf{p}} dV_p \mathcal{P} = \int_{R, \mu, v} dV_p \mathcal{P} = \frac{2^{2+2\ell_l} E_{nl}^{3+2\ell_l}}{(1+\ell_l)^{4+2\ell_l} \Gamma(2+2\ell_l)} r^{2\ell_l} e^{-2rE_{nl}/(1+\ell_l)} \quad (2.20)$$

On the other hand, the spatial average yields

$$dV_p \int_0^\infty dr r^2 \mathcal{P} = dR \frac{d\mu}{\pi} \frac{dv}{2\pi} \frac{E_{nl}^{5+4\ell_l}}{(1+\ell_l)^{6+4\ell_l} \Gamma(\frac{3}{2}+2\ell_l) \Gamma(\frac{5}{2}+2\ell_l)} R^{4+4\ell_l} \sin^{2+4\ell_l} \mu e^{-2RE_{nl}/(1+\ell_l)} \quad (2.21)$$

The angular integral brings

$$\int_0^\infty dr r^2 \int_{\mu, v} dV_p \mathcal{P} = dR \frac{2^{3+4\ell_l} E_{nl}^{5+4\ell_l}}{(1+\ell_l)^{6+4\ell_l} \Gamma(4+4\ell_l)} R^{4+4\ell_l} e^{-2RE_{nl}/(1+\ell_l)} \quad (2.22)$$

This implies the expectation values

$$\langle \frac{1}{r} \rangle = \frac{E_{nl}}{(1+\ell_l)^2} = \langle \frac{2}{R} \rangle, \quad \langle \frac{1}{r^2} \rangle = \frac{2E_{nl}^2}{(1+\ell_l)^3(1+2\ell_l)}, \quad \langle \frac{1}{R^2} \rangle = \frac{E_{nl}^2}{(1+\ell_l)^3(3+4\ell_l)} \quad (2.23)$$

Eq. (2.5) then implies, to order  $\alpha^4$ , the classical expectation value

$$\langle E \rangle_{nl; cl} = 1 - \alpha^2 Z^2 \langle \frac{1}{R} \rangle + \frac{1}{2} \alpha^4 Z^4 \langle \frac{1}{R^2} \rangle = 1 - \frac{\alpha^2 Z^2}{2} - \frac{(4n^2 + 4n - 1)\alpha^4 Z^4}{4n^4(8n^2 - 6n + 1)} = E_{nl} + \frac{\alpha^4 Z^4}{8n^4(4n - 1)} \quad (2.24)$$

This is not the quantum mechanical result, so the approach fails at order  $\alpha^4$ , that is to say, for the fine structure.

## NEGLECTING CORRELATIONS

It is easily verified that the proper  $\alpha^4$  term of the quantum eigenenergies (1.13) emerges if we replace the term  $\langle R^{-2} \rangle$  of Eq. (2.24) by  $\langle R^{-1} \rangle^2$ , thus neglecting the correlator  $\langle R^{-2} \rangle - \langle R^{-1} \rangle^2$ . This approach appears to work to infinite order.

## Yrast states

Surprisingly, the quantum mechanical result does emerge if, after expanding the square root (2.5) in powers of  $\alpha$ , we replace in the successive terms the averages of products by products of averages. The result can then be resummed,

$$\langle \langle E \rangle \rangle_{nl; cl} \equiv \sqrt{1 + \alpha^4 Z^4 \langle \frac{1}{R} \rangle^2} - \alpha^2 Z^2 \langle \frac{1}{R} \rangle = \sqrt{1 + \frac{\alpha^4 Z^4 E_{nl}^2}{4(1+\ell_l)^4}} - \frac{\alpha^2 Z^2 E_{nl}}{2(1+\ell_l)^2} \quad (3.1)$$

From the expression (1.13) one may derive the relation

$$\alpha^2 Z^2 = \frac{1 - E_{nl}^2}{E_{nl}^2} (n - l + \ell_l)^2 \quad (3.2)$$

For the Yrast states we then observe agreement with the quantum eigenvalues for the energy,

$$\langle \langle E \rangle \rangle_{nl; cl} = \sqrt{\frac{(1+E_{nl}^2)^2}{4E_{nl}^2} - \frac{1-E_{nl}^2}{2E_{nl}}} = E_{nl} \quad (3.3)$$

to all orders in  $\alpha$ . Physically such infinite precision is fortuitous, since there are Lamb shift corrections of order  $\alpha^5$ .

By working with expression (2.4), we have another approach, which acts as a consistency test. If we separate the averages and calculate the average of  $E$  as above, we find

$$\langle \langle p_r^2 \rangle \rangle = \langle \langle 2E(\frac{1}{r} - \frac{1}{R}) \cos^2 \mu \rangle \rangle = 2\langle \langle E \rangle \rangle \langle (\frac{1}{r} - \frac{1}{R}) \cos^2 \mu \rangle = \frac{E_{nl}^2}{4(1+\ell_l)^3} \quad (3.4)$$

and

$$\langle \langle \frac{\omega^2 L^2}{r^2} \rangle \rangle = \frac{E_{nl}^2 (3+4\ell_l)}{4(1+\ell_l)^3}. \quad (3.5)$$

These results bring

$$\langle \langle \frac{L^2}{r^2} \rangle \rangle = \langle \langle \frac{\omega^2 L^2}{r^2} \rangle \rangle + \langle \langle \frac{\alpha^2 Z^2}{r^2} \rangle \rangle = \langle \langle \frac{\omega^2 L^2}{r^2} \rangle \rangle + \alpha^2 Z^2 \langle \frac{1}{r} \rangle^2 = \frac{1}{(1+\ell_l)^2} - \frac{E_{nl}^2}{4(1+\ell_l)^3} \quad (3.6)$$

Combining both results brings simply

$$\langle\langle \mathbf{p}^2 \rangle\rangle = \langle\langle p_r^2 \rangle\rangle + \langle\langle \frac{L^2}{r^2} \rangle\rangle = \frac{1}{(1+\ell_l)^2} \quad (3.7)$$

When averaging Eq. (2.4), we again expand in powers of  $\alpha^2 Z^2$  and replace repeatedly averages of products by products of averages. We end up with the alternative expression

$$\langle\langle E \rangle\rangle'_{nl;cl} = \sqrt{1 + \alpha^2 Z^2 \langle\langle \mathbf{p}^2 \rangle\rangle} - \langle\langle \frac{\alpha^2 Z^2}{r} \rangle\rangle = \sqrt{1 + \frac{\alpha^2 Z^2}{(1+\ell_l)^2}} - \frac{\alpha^2 Z^2 E_{nl}}{(1+\ell_l)^2} \quad (3.8)$$

The leading correction of this square root is stronger than the one of order  $\alpha^4 Z^4$  of the expression (3.1). But by using eq. (3.2) is easy to see that

$$\langle\langle E \rangle\rangle'_{nl;cl} = \sqrt{\frac{1}{E_{nl}^2} - \frac{1 - E_{nl}^2}{E_{nl}}} = E_{nl} \quad (3.9)$$

confirming consistency of the approach, provided it includes the replacement  $\langle\langle r^{-2} \rangle\rangle \rightarrow \langle r^{-1} \rangle^2$  done in Eq. (3.6).

## The 2s state

The Yrast states could be special, because their spatial wavefunctions consist of a single term. We now investigate a non-Yrast state, the 2s state, which has an additional polynomial in its wavefunction:

$$R_{20} = \left\{ \frac{(1+\ell_0)2^{3+2\ell_0}}{(2+\ell_0)(4+3\ell_0)^{3/2+\ell_0}\Gamma(2+2\ell_0)} \right\}^{1/2} r^{\ell_0} \left( 1 - \frac{r}{(1+\ell_0)\sqrt{4+3\ell_0}} \right) \exp\left(-\frac{r}{\sqrt{4+3\ell_0}}\right) \quad (3.10)$$

The function  $R_{20}(\frac{r}{E_{20}})$  is orthogonal to  $R_{10}(\frac{r}{E_{10}})$ , as it should in view of the scaled Schrödinger equation (1.9). The probability density  $P_{20}(r) = R_{20}^2(r)$  may derive from a phase space density,

$$\begin{aligned} \mathcal{P}_{20}(E, L, R(E)) &= \frac{2^{7+8\ell_0}}{(2+\ell_0)\pi^2\Gamma(6+4\ell_0)(4+3\ell_0)^{9/2+2\ell_0}} \omega L R^3 \Phi(E) \left( \frac{\omega^2 L^2 R}{2E} \right)^{2\ell_0} \exp\left(-\frac{2R}{\sqrt{4+3\ell_0}}\right) \\ &\times \left[ (1+\ell_0)^2 (4+3\ell_0)^2 (3+4\ell_0) (5+4\ell_0) - 8(1+\ell_0)(4+3\ell_0)(5+4\ell_0) \frac{\omega^2 L^2 R}{2E} + 16 \left( \frac{\omega^2 L^2 R}{2E} \right)^2 \right] \end{aligned} \quad (3.11)$$

which after integrating over the angles  $\mu$  and  $\nu$  reads

$$\begin{aligned} \int_{\mu, \nu} dV_p \mathcal{P}_{20} &= \frac{dR^{28+4\ell_0}}{(4+3\ell_0)^{9/2+2\ell_0}\Gamma(3+2\ell_0)\Gamma(5+2\ell_0)} r^{2\ell_0} (R-r)^{1+2\ell_0} \exp\left(-\frac{2R}{\sqrt{4+3\ell_0}}\right) \\ &\times [(1+\ell_0)^3 (3+2\ell_0) (4+3\ell_0)^2 - 2(1+\ell_0)(3+2\ell_0)(4+3\ell_0)r(R-r) + 2r^2(R-r)^2] \end{aligned} \quad (3.12)$$

and next

$$\begin{aligned} \int_0^R dr r^2 \int_{\mu, \nu} dV_p \mathcal{P}_{20} &= \frac{2^{9+4\ell_0} (3+2\ell_0)}{(4+3\ell_0)^{9/2+2\ell_0}\Gamma(9+4\ell_0)} R^{4+4\ell_0} \exp\left(-\frac{2R}{\sqrt{4+3\ell_0}}\right) \\ &\times [(1+\ell_0)^2 (4+3\ell_0)^2 (5+4\ell_0) (7+4\ell_0) - 2(1+\ell_0)(4+3\ell_0)(7+4\ell_0)R^2 + R^4] \end{aligned} \quad (3.13)$$

This brings the averages

$$\langle \frac{1}{r} \rangle = \frac{1}{(2+\ell_0)\sqrt{4+3\ell_0}}, \quad \langle \frac{1}{R} \rangle = \frac{1}{2(2+\ell_0)\sqrt{4+3\ell_0}}, \quad (3.14)$$

which is just what is needed, because we get using  $\alpha^2 Z^2 = -\ell_0(\ell_0 + 1)$ ,

$$\langle\langle E \rangle\rangle_{20;cl} = \sqrt{1 + \alpha^4 Z^4 \langle \frac{1}{R} \rangle^2} - \langle \frac{\alpha^2 Z^2}{R} \rangle = \frac{8 + 7\ell_0 + \ell_0^2}{2(2 + \ell_0)\sqrt{4 + 3\ell_0}} + \frac{\ell_0(\ell_0 + 1)}{2(2 + \ell_0)\sqrt{4 + 3\ell_0}} = \frac{2 + \ell_0}{\sqrt{4 + 3\ell_0}} = E_{20} \quad (3.15)$$

We can further verify that

$$\langle\langle \mathbf{p}^2 \rangle\rangle = 2\langle\langle E \rangle\rangle \left[ \langle \frac{1}{r} \rangle - \langle \frac{1}{R} \rangle \right] + \alpha^2 Z^2 \langle \frac{1}{r} \rangle^2 = \frac{1}{(2 + \ell_0)^2} \quad (3.16)$$

implying that also the alternative calculation leads to the quantum mechanics result,

$$\langle\langle E \rangle\rangle'_{20;cl} = \sqrt{1 + \alpha^2 Z^2 \langle\langle \mathbf{p}^2 \rangle\rangle} - \alpha^2 Z^2 \langle \frac{1}{r} \rangle = \frac{\sqrt{4 + 3\ell_0}}{2 + \ell_0} + \frac{\ell_0(\ell_0 + 1)}{(2 + \ell_0)\sqrt{4 + 3\ell_0}} = \frac{2 + \ell_0}{\sqrt{4 + 3\ell_0}} \quad (3.17)$$

It appears that the approach is non-unique. The density  $R_{20}^2(r)$  can also be obtained from

$$\begin{aligned} \mathcal{P}'_{20} &= \frac{2^{6+4\ell_0}(1+\ell_0)}{(2+\ell_0)\Gamma^2(3+2\ell_0)(4+3\ell_0)^{7/2+2\ell_0}} \omega L R^3 \Phi(E) \left( \frac{\omega^2 L^2 R}{2E} \right)^{2\ell_0} \exp\left(-\frac{2R}{\sqrt{4+3\ell_0}}\right) \\ &\times \left[ (1+\ell_0)(4+3\ell_0)(7+8\ell_0) - 8(1+\ell_0)\sqrt{4+3\ell_0}R + 2R^2 \right] \left[ 2\pi B \left( \frac{3}{2} + 2\ell_0, \frac{1}{2} \right) \right]^{-1} \end{aligned} \quad (3.18)$$

This form has the same expectation for  $\langle 1/r \rangle$  and  $\langle 1/R \rangle$ , and thus for energy. The two forms can thus be mixed, which leads to a one-parameter family of solutions; the proper choice cannot be fixed in the present approach. By viewing such mixings as a way to fix the  $r^2$  term of the square of the polynomial of  $R_{20}$ , one might wonder whether the  $r$  term can be changed with a second independent parameter. It has been checked, however, that that would modify  $\langle 1/R \rangle$ .

## PHASE SPACE FORMS FOR SQUARES OF SPHERICAL HARMONICS

Previous expressions  $R_{n0}^2$  for  $s = 0$  states are still multiplied by a factor  $Y_{00}^2 = 1/(4\pi)$ , which drops out after the integration over the angles  $\theta$  and  $\phi$  of the coordinate vector. Let us denote the generalization of  $Y_{lm}^* Y_{l'm'}$  by  $\mathcal{Y}_{lm}^{l'm'}$ . Then we can likewise multiply our results for  $\mathcal{P}_{10}$  and  $\mathcal{P}_{20}$  by  $\mathcal{Y}_{00}^{00} = 1/(4\pi)$  and cancel it by the angular integrals.

For the  $\mathcal{P}_{21}$  distribution we only considered the radial part; there is still the question how to deal with the factor  $Y_{1m}^2$  and the same question arises for the higher  $l$  states at  $n \geq 2$ . According to the quantum solution, such diagonal terms should have no time-dependence, so also  $Y_{1m}^2$  is be generalized to a function depending on conserved quantities only.

Let us notice that in the frame along  $\mathbf{r}$  we can introduce angles  $\mu$  and  $\nu$  with  $0 \leq \mu \leq \pi$  and  $0 \leq \nu \leq 2\pi$ ,

$$\mathbf{r} = r(0, 0, 1), \quad \mathbf{p} = p(\sin \mu \cos \nu, \sin \mu \sin \nu, \cos \mu) = (p_{\perp 1}, p_{\perp 2}, p_r), \quad \mathbf{L} = pr \sin \mu (-\sin \nu, \cos \nu, 0) \quad (4.1)$$

In the laboratory frame this implies

$$\begin{aligned} \mathbf{r} &= r(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta), \quad \mathbf{L} = pr \sin \mu \hat{\mathbf{L}}, \\ \hat{\mathbf{L}} &= (-\cos \theta \cos \phi \sin \nu - \sin \phi \cos \nu, -\cos \theta \sin \phi \sin \nu + \cos \phi \cos \nu, \sin \theta \sin \nu) \end{aligned} \quad (4.2)$$

Specifying the  $z$ -axis as our preferred axis, we have  $\hat{L}_z = \sin \theta \sin \nu$  and we define

$$\mathcal{Y}_{1-1}^{1-1} = \frac{3}{4\pi} (\hat{L}_z^2 - \hat{L}_z), \quad \mathcal{Y}_{10}^{10} = \frac{3}{4\pi} (1 - 2\hat{L}_z^2), \quad \mathcal{Y}_{11}^{11} = \frac{3}{4\pi} (\hat{L}_z^2 + \hat{L}_z). \quad (4.3)$$

In general with  $\mathbf{n}$  a fixed unit vector this would read

$$\mathcal{Y}_{1-1}^{1-1} = \frac{3}{4\pi} [(\mathbf{n} \cdot \hat{\mathbf{L}})^2 - \mathbf{n} \cdot \hat{\mathbf{L}}], \quad \mathcal{Y}_{10}^{10} = \frac{3}{4\pi} [1 - 2(\mathbf{n} \cdot \hat{\mathbf{L}})^2], \quad \mathcal{Y}_{11}^{11} = \frac{3}{4\pi} [(\mathbf{n} \cdot \hat{\mathbf{L}})^2 + \mathbf{n} \cdot \hat{\mathbf{L}}], \quad (4.4)$$

The odd terms are chosen linear, as this seems most natural to us. To settle their form uniquely is, however, not possible in the present approach, since for that aim the underlying physics should be specified further.

The  $\mathbf{p}$ -integration produces an integral over  $v$ , which brings the desired results

$$\int_0^{2\pi} \frac{dv}{2\pi} \mathcal{Y}_{11}^{11} = \int_0^{2\pi} \frac{dv}{2\pi} \mathcal{Y}_{1-1}^{1-1} = \frac{3}{8\pi} \sin^2 \theta = |Y_{1\pm 1}|^2, \quad \int_0^{2\pi} \frac{dv}{2\pi} \mathcal{Y}_{10}^{10} = \frac{3}{4\pi} \cos^2 \theta = |Y_{10}|^2 \quad (4.5)$$

Moreover, it holds that

$$\int_0^{2\pi} \frac{dv}{2\pi} \hat{L}_z \mathcal{Y}_{1m}^{1m} = \frac{3m}{8\pi} \sin^2 \theta, \quad (m = -1, 0, 1) \quad (4.6)$$

If we integrate this over all angles  $\theta, \phi$  and make the identification  $L_{zop} \leftrightarrow \hat{L}_z$ , we reproduce the quantum result

$$\overline{\hat{L}_z \mathcal{Y}_{1m;1m}} \equiv \int \sin \theta d\theta d\phi \int \frac{dv}{2\pi} \hat{L}_z \mathcal{Y}_{1m}^{1m} = m = \langle Y_{1m} | L_{zop} | Y_{1m} \rangle, \quad (m = -1, 0, 1) \quad (4.7)$$

and also the related results

$$\overline{\hat{L}_x \mathcal{Y}_{1m}^{1m}} = 0 = \langle Y_{1m} | L_{xop} | Y_{1m} \rangle, \quad \overline{\hat{L}_y \mathcal{Y}_{1m}^{1m}} = 0 = \langle Y_{1m} | L_{yop} | Y_{1m} \rangle. \quad (4.8)$$

We conclude that at least the lowest non-trivial square of spherical harmonics can be properly generalized.

## COMPARISON WITH THE WIGNER FUNCTION

The Wigner function of the ground state wavefunction reads in the non-relativistic limit [24, 25]

$$W(\mathbf{r}, \mathbf{p}) = \int d^3q \psi^*(\mathbf{p} + \frac{1}{2}\mathbf{q}) \psi(\mathbf{p} - \frac{1}{2}\mathbf{q}) e^{-i\mathbf{r} \cdot \mathbf{q}} = \frac{2^6 \pi}{(2\pi)^6} \int d^3q \frac{e^{-i\mathbf{r} \cdot \mathbf{q}}}{[1 + (\mathbf{p} + \frac{1}{2}\mathbf{q})^2]^2 [1 + (\mathbf{p} - \frac{1}{2}\mathbf{q})^2]^2} \quad (5.1)$$

As opposed to our phase space density, this function is partly negative, a fact commonly considered as a key argument against any underlying quasi-classical reality. Its marginal reads

$$W(\mathbf{p}) = \frac{8}{\pi^2 (1 + p^2)^4} \quad (5.2)$$

The non-relativistic limit of our phase space density (2.17) reads for the ground state

$$\mathcal{P}(\mathbf{r}, \mathbf{p}) = \frac{2}{\pi^3} LR^3 e^{-2R} \quad (5.3)$$

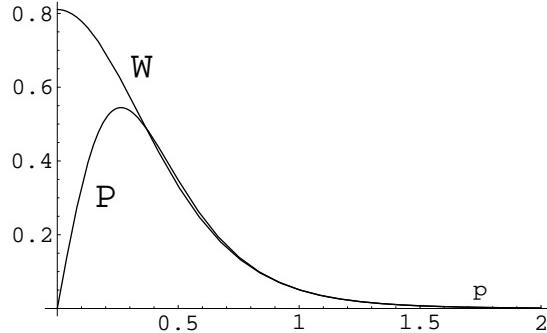
Taking spherical coordinates along  $\mathbf{p}$  yields  $L = pr \sin \theta$  and using  $\frac{1}{2}p^2 - 1/r = -1/R$  one derives for its marginal

$$\mathcal{P}(\mathbf{p}) = \frac{2p}{\pi} \int_0^{2/p^2} dr r^3 R^3 e^{-2R} = \frac{2p}{\pi} \int_0^\infty dR \frac{R^6}{(1 + \frac{1}{2}p^2 R)^5} e^{-2R} \quad (5.4)$$

Whereas (5.2) is finite at  $p = 0$  and decays as  $p^{-8}$  for large  $p$ , Eq. (5.4) is linear at  $p = 0$  and decays as  $p^{-9}$  for large  $p$ . Figure 5.1 compares these results. It is seen that the difference is mainly at small  $p$ , where because of the phase space volume little weight is located. The tails are already small when they start to deviate from each other.

## DISCUSSION

In this paper we have considered the question whether the quantum probability densities (the squares of quantum wavefunctions) of the relativistic hydrogen atom, in the absence of electron spin, may arise from the phase space distribution in an underlying classical stochastic theory (“sub-quantum theory” or “hidden variable theory”). Because of the small value of the fine structure constant, this is a weak damping problem, for which it is generally known that the stationary distribution is determined by conserved quantities (energy and angular momentum) alone. Without



**FIGURE 5.1.** Reduced phase space density  $\mathcal{P}(p)$  of the non-relativistic hydrogen ground state, Eq. (5.4), (lower curve), and reduced Wigner density  $W(p)$ , Eq. (5.2), (upper curve), as function of the momentum  $p$ , in atomic units. They are both normalized to unity.

specifying the underlying stochastic theory, we have shown by construction that such a scheme can work for the considered cases. The relativistic version of the hydrogen problem has been chosen in order to have the fine structure constant as a parameter, and the generalization works to all orders (be it that from the outset, all Lamb shift effects are neglected, in our approach as well as in Schrödinger dynamics).

Our resulting phase space density  $\mathcal{P}(\mathbf{r}, \mathbf{p})$  does not coincide with the Wigner function  $W(\mathbf{r}, \mathbf{p})$ . For the hydrogen ground state the latter can be expressed as derivatives of an integral [24, 25]. In contrast, for our ground state  $\mathcal{P}(\mathbf{r}, \mathbf{p})$  is explicit, it contains powers and an exponential. Although we did not prove its uniqueness, it looks simple and has been generalized for the whole family of Yrast states. It is well known that the Wigner function for the hydrogen ground state is partly negative, and this has long been considered as a no-go argument against any classical-type subquantum structure. However, our  $\mathcal{P}(\mathbf{r}, \mathbf{p})$  is non-negative for the ground state. Does this solve that paradox? Yes, because our momentum stands for the instantaneous momentum of the particle. The momentum of the Wigner function is defined in via off-diagonal elements of the quantum density matrix, and should be considered as a coarse grained momentum, just as would happen in brownian motion [26, 11]. For such a coarse grained momentum positivity of the density is not obvious, probably not needed, and certainly absent.

From a physical point of view, there are many reasons to accept that particles are to good approximation point particles: traces in cloud chambers, images of non-flat crystal surfaces obtained by scanning tunnel microscopy, trapping of ions in magnetic fields over periods exceeding a year. If, one the subquantum level they just have a position and a speed, the phase space density cannot be negative. The problem of negative probabilities is solved for our ground state, but still a potential danger for excited states. For the  $2p$  state the generalized spherical harmonics are partly negative, while for the  $2s$  state the radial phase space density is partly negative. But in the theory of classical brownian motion the excited states are also partly negative, while the total probability density will always be non-negative.

Within Stochastic Electrodynamics the  $2s$  state can be populated by a two-photon excitation from an ensemble of ground state atoms, but it cannot happen in such a manner that the total phase space probability is negative anywhere. For instance, one could imagine a large population of the  $1s$  state and a partial one of the  $2s$  state, with not-too-large values of  $L$  and  $R$  in the  $2s$  state, such that the total probability is still positive everywhere. Afterall, at the classical level each realization consists of an electron in some orbit around the nucleus in the presence of fluctuations, and a ‘state’ merely describes an ensemble of such systems. The  $2p$  phase space densities at  $m$  fixed are partly negative, so they cannot exist on their own. However, the ‘transversally’ polarized mixture  $\mathcal{P}_{21}(E, L)(\mathcal{Y}_{1-1}^{1-1} + \mathcal{Y}_{11}^{11})$  is non-negative, while the unpolarized mixture  $\mathcal{P}_{21}(E, L)(\mathcal{Y}_{1-1}^{1-1} + \mathcal{Y}_{10}^{10} + \mathcal{Y}_{11}^{11}) = 3\mathcal{P}_{21}(E, L)/(4\pi)$  is strictly positive. They can in principle exist without ground state buffer. The odd parts may differ from the assumed linear ones of (4.4). When they are cubic, both  $\mathcal{Y}_{1-1}^{1-1}$  and  $\mathcal{Y}_{11}^{11}$  can be positive, but not  $\mathcal{Y}_{10}^{10}$ . The quantum operator  $L_{zop}$  may then be identified with the classical  $\hat{L}_z$  times a constant.

Given the phase space distribution, we have considered the average energy. We have observed that the quantum energy eigenvalues are reproduced provided certain correlations are neglected, after which averages of products reduce to products of averages. Physically this occurs when these averages are taken at very different times, such that their fluctuations are uncorrelated. This hints at the possibility that the operator structure of quantum theory arises from the physics of timescale separated observables in an underlying stochastic theory. The property that repeatedly applying a quantum operator on one of its eigenstates produces just powers of the related eigenvalue has a natural interpretation

in terms of temporal separation of averages, as is long understood in the theory Stochastic Electrodynamics [9]. It is remarkable that such a separation appears necessary here to establish the proper mapping of energies between the class of stochastic theories and quantum mechanics.

In atomic physics spectral emissions are due to transition from an excited state to a lower one. The  $2p$ -state has a lifetime given by  $\alpha^{-3}\tau_0$  multiplied by phase space factors, the final result is of order of magnitude of the inverse Lamb shift frequency  $\sim 1\text{ ns}$ . So about  $\alpha^{-3}$  Keppler orbits are traversed before the emission takes place. This large number of revolutions, together with the gradual influence of the noise that slowly changes the orbits, may explain the sharpness of spectral lines within our picture. If only a few orbits are traversed, the sharpness of the spectral line will not be achieved, since the spread in energies of Keppler orbits is large. But this just provides a physical picture for the energy-time uncertainty relation known from quantum mechanics.

There are many remaining and often worrying issues. An essential item is the role of dynamics, such as the  $1s$ - $2s$  and  $1s$ - $2p$  interference terms. Likewise, there is the task to include spin. It is good to realize that our present radial distributions are functions of  $\mathbf{L}^2$ , which remains conserved in the presence of spin.

For scattering and ionization experiments on e.g. electrons on e.g. atomic hydrogen [27] it is known that the energy loss of the incoming electron can only take one of the discrete values needed to bring the ground state electron to the one of the excited levels (Lyman lines). In the recent work [28] transitions to  $n = 2, 3$  and  $4$  are clearly observed, while the higher states form a quasi continuum, that matches the ionization continuum. Whether these discrete features can be described in the present formalism is a question related to the existence of excited states, and the timescale on which they express themselves. What we have discussed is a number of features that should be respected in our class of the underlying theories. Whether the proposed type of ground state and excited states exists and what happens during electron scattering events, are questions that cannot be answered before specializing the precise type of stochastic noise.

Like any new theory, the present approach poses more questions than it solves. For any sub-quantum theory a vast amount of experimental constraints is to be respected. But we are convinced that no individual quantum experiment is described by any current theory. So if Einstein was right after all, quantum mechanics is incomplete and then there should be some path to an underlying, more complete theory. We hope that the present work plays a role in uncovering it.

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